



THE UNIVERSITY OF
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Properties of Ionospheric Doppler Oscillations Driven by Downgoing ULF Waves

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Background

ULF wave signatures are often observed in SD ground scatter returns at mid- to low latitudes.

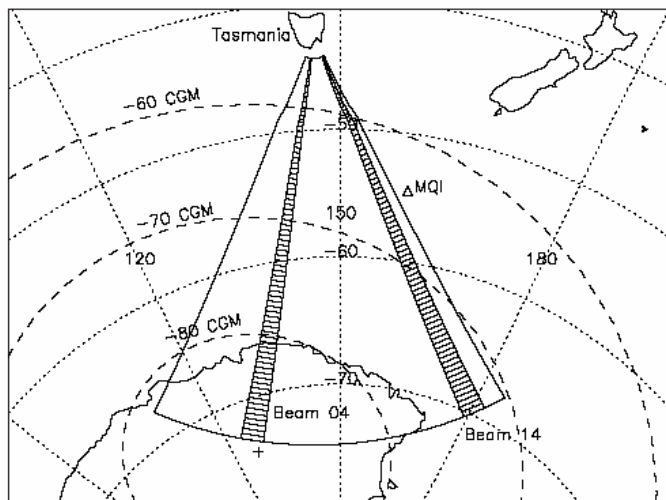
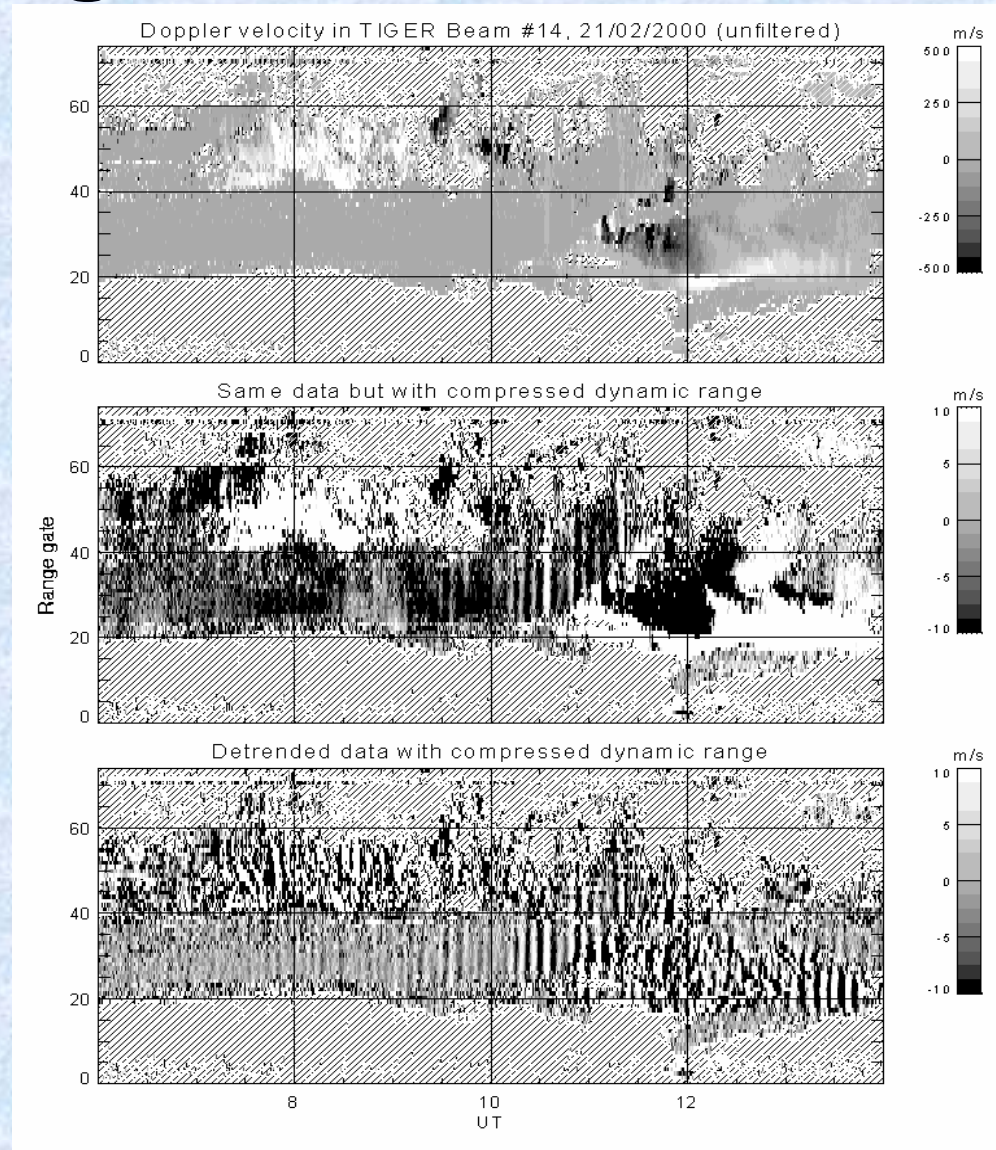


Figure 1. TIGER FOV in geographic coordinates. Triangle denotes Macquarie Island (MQI) magnetometer site.



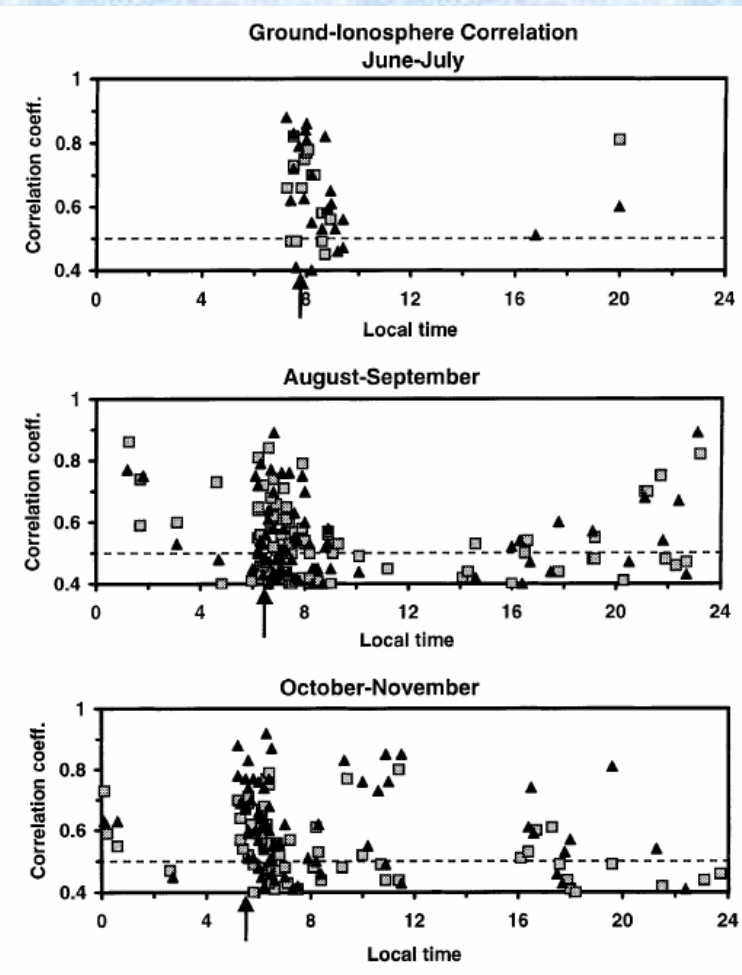
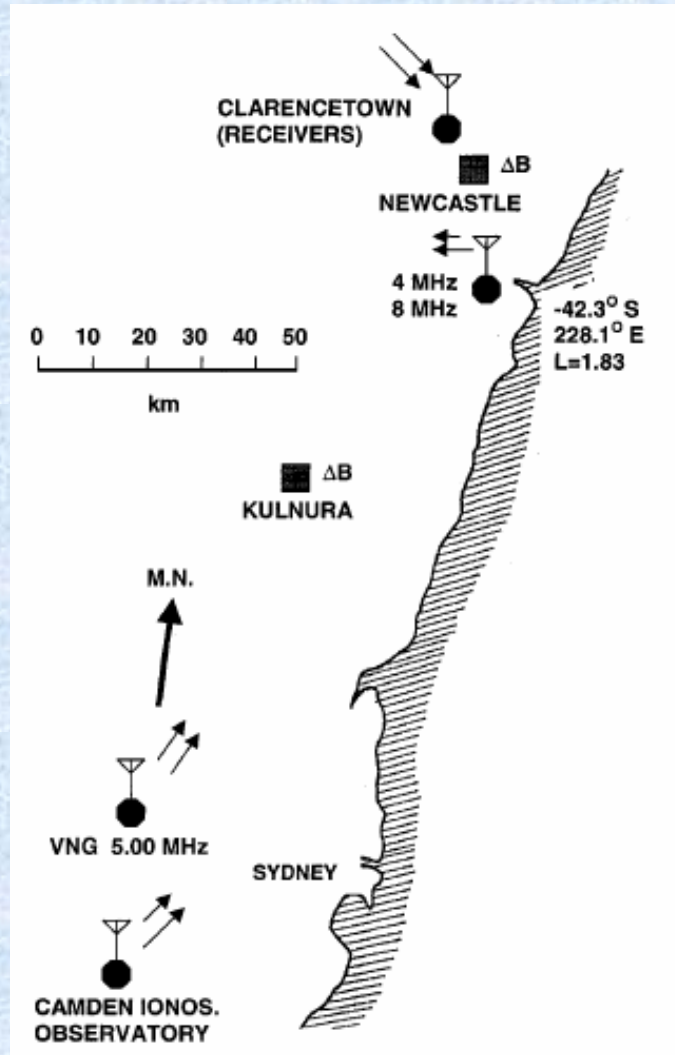
[Ponomarenko, Menk and Waters, *GRL*, **30**, 2003]

The Experiments

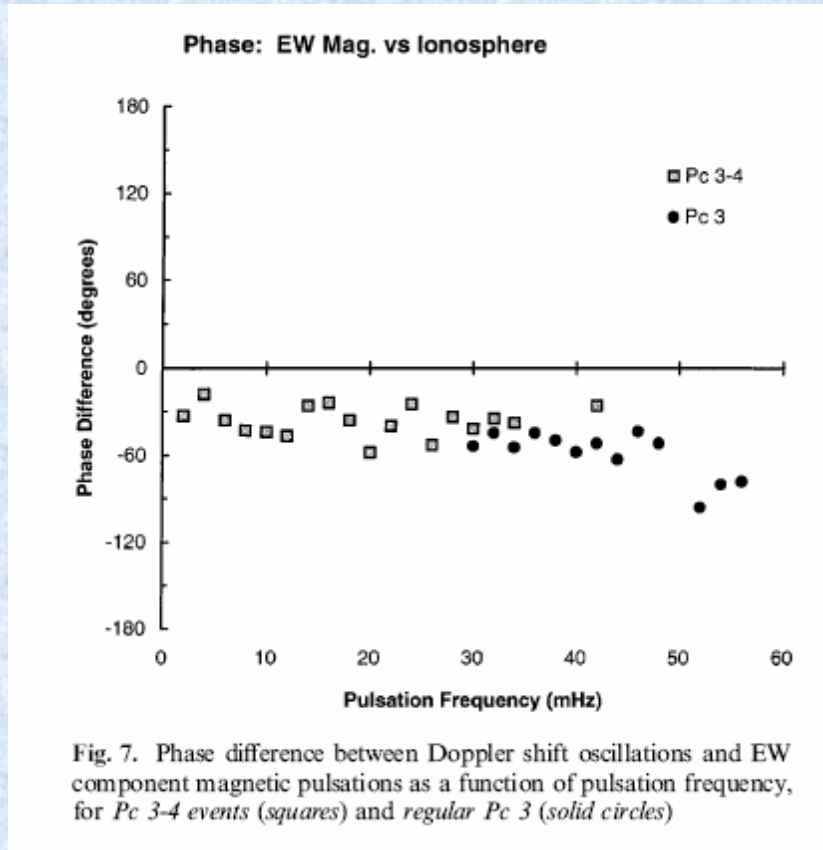
- We investigated the properties of ULF wave fields in the low-latitude ionosphere using HF Doppler radio sounders and ground magnetometers.
- The Doppler sounders monitored frequency shifts in O-mode reflections from HF frequency standard transmitters with a sensitivity of ≤ 0.005 Hz over 1-140 mHz.
- The magnetometers provided information on the ULF wave modes propagating from the magnetosphere through the ionosphere.
- There were two studies:
 - A preliminary study at $L = 1.83$
 - A detailed study spanning $L = 1.57 - 2.77$.

Preliminary Study

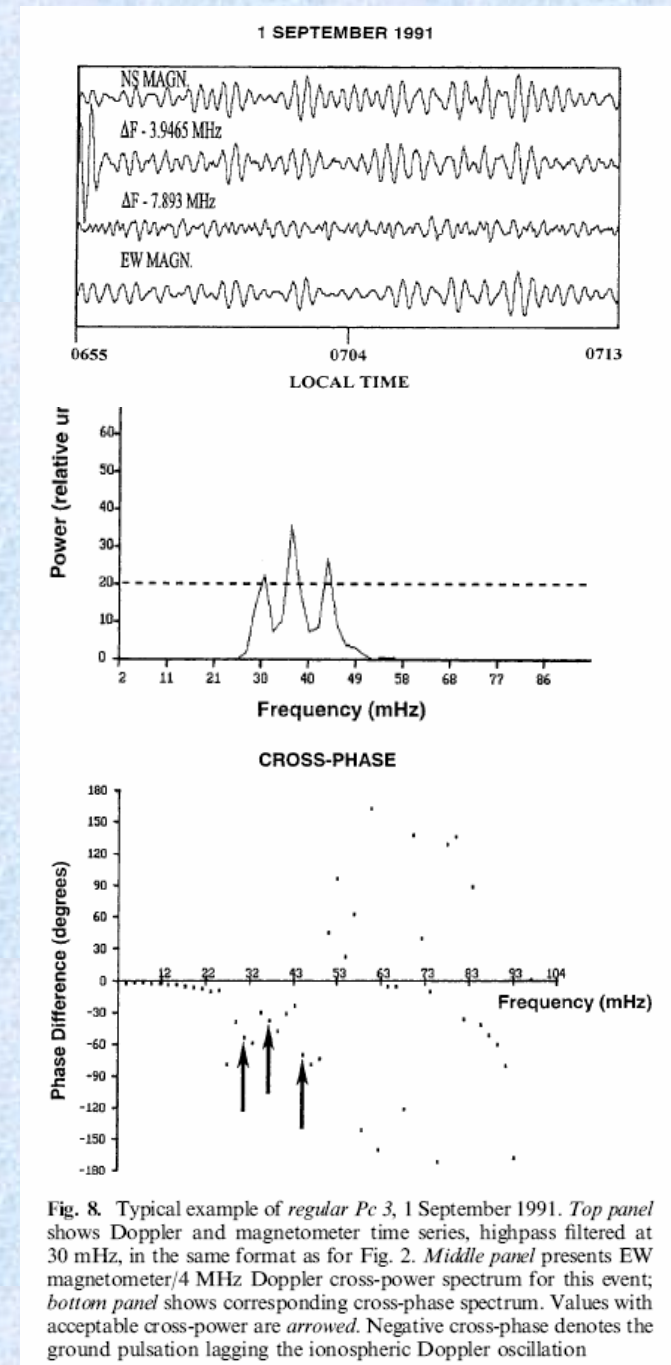
$L = 1.83$; Jun – Dec 1991
and Dec 1992 – Jul 1993.



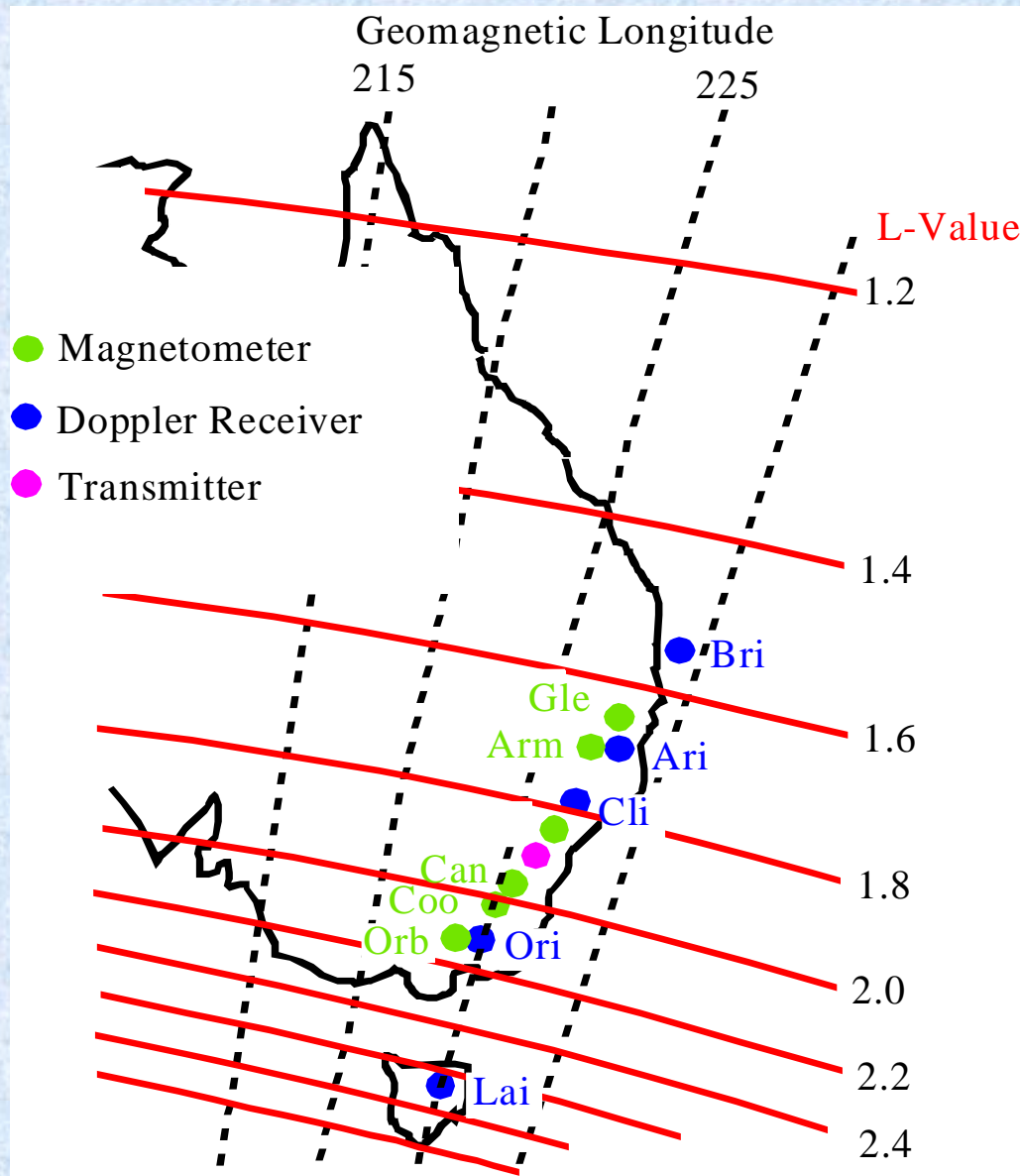
Preliminary Study



[Marshall and Menk,
Ann. Geophys., 17, 1999]



Detailed Study



HF transmitters at $L = 1.94$.

5 Doppler receivers

10 magnetometers arranged in 5 pairs beneath ionospheric midpoints.

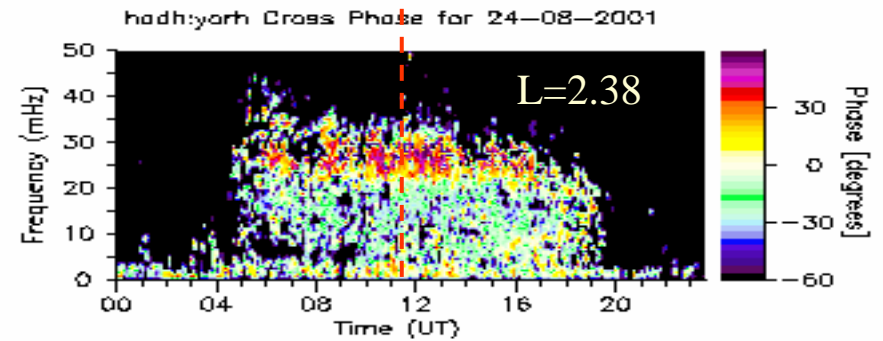
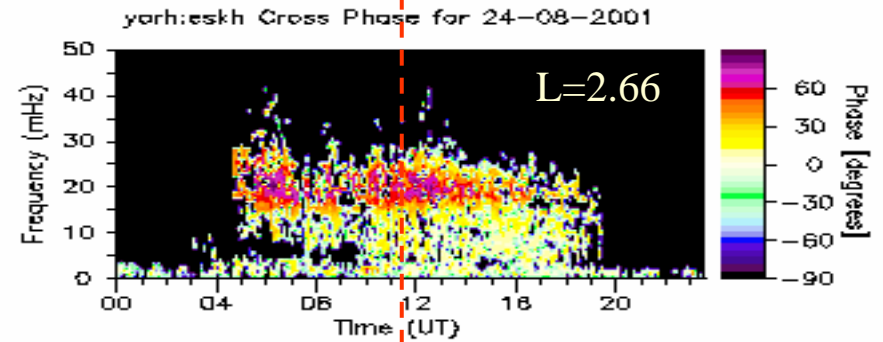
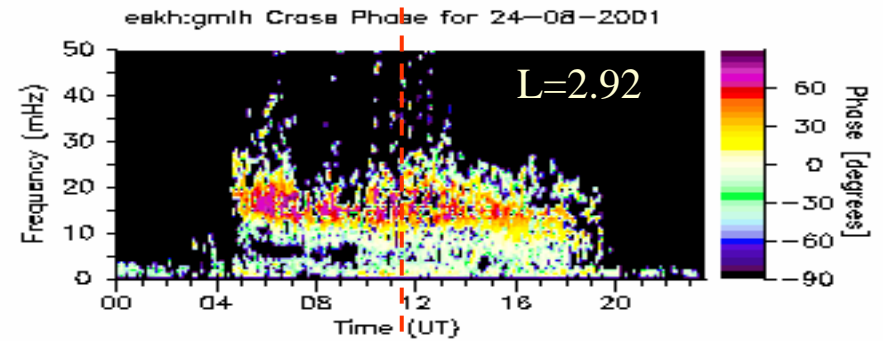
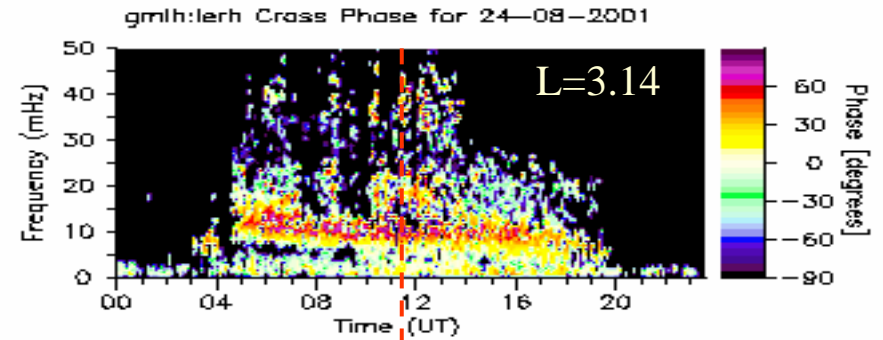
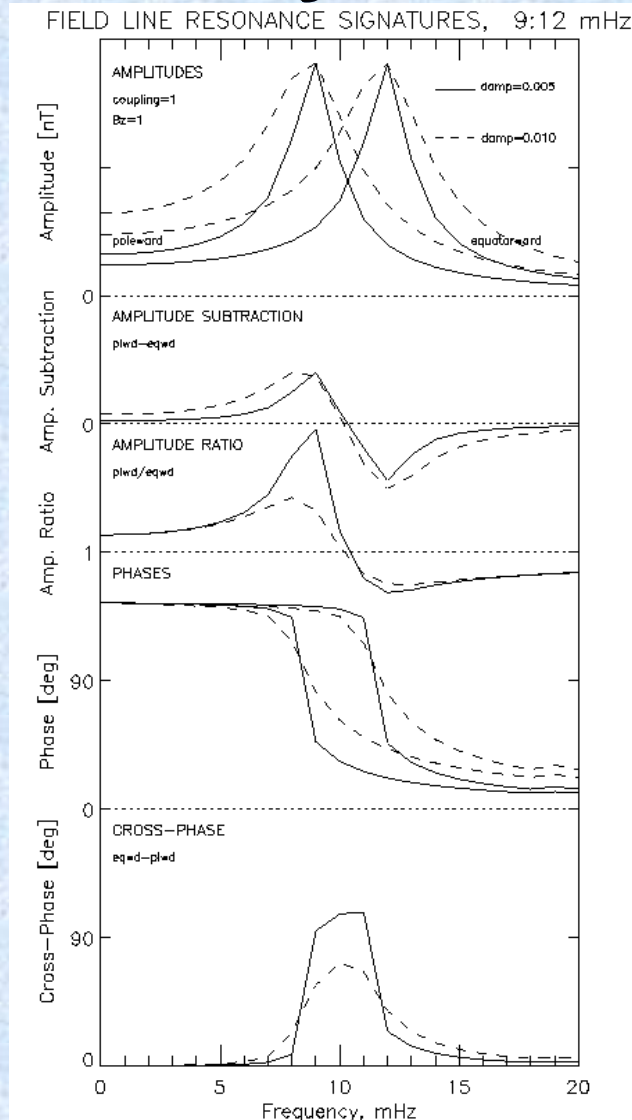
Dec 1992 – Jul 1993
and Jan – April 1994.

[Menk et al., *GRL*, **34**,
2007]

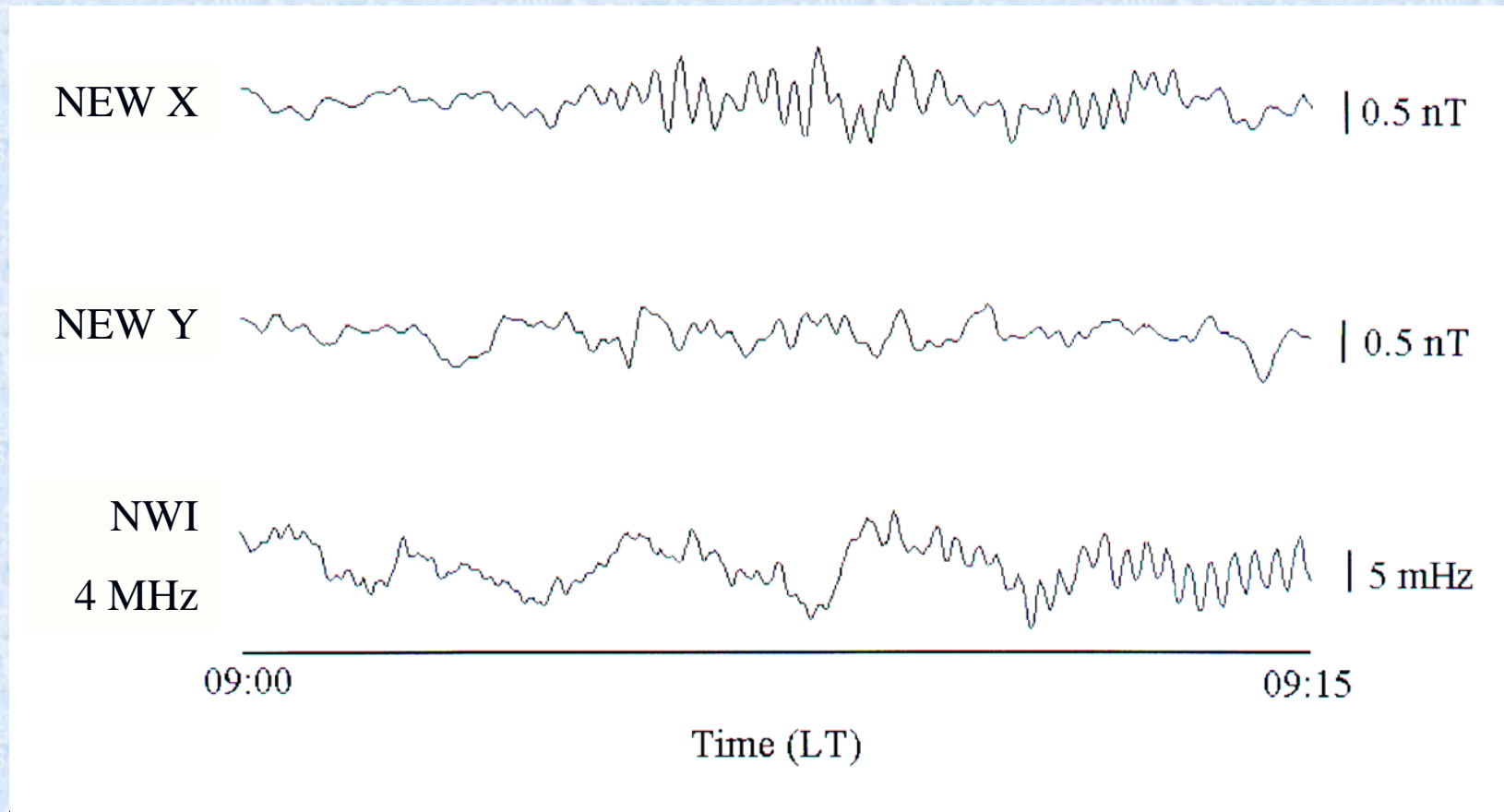
Data Analysis

- Radio frequencies chosen gave similar F-region reflection height across the array.
- Examined time series, power and phase spectra of ULF pulsations in the ionosphere and on the ground.
- Ionosphere-ground power ratio and cross-phase were determined as a function of ULF frequency.
- Pure state vector filtering used on the Doppler data, set to a polarization state best representing azimuth and ellipticity over 25 days of noise-free data in 1993.

Magnetometer Analysis

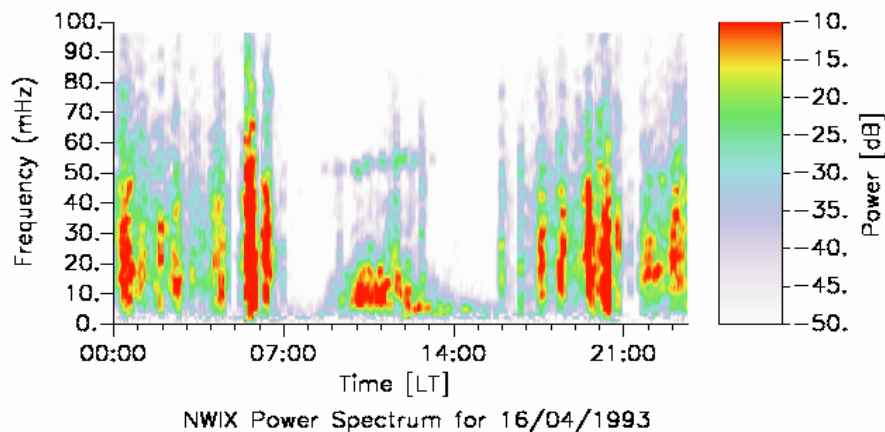
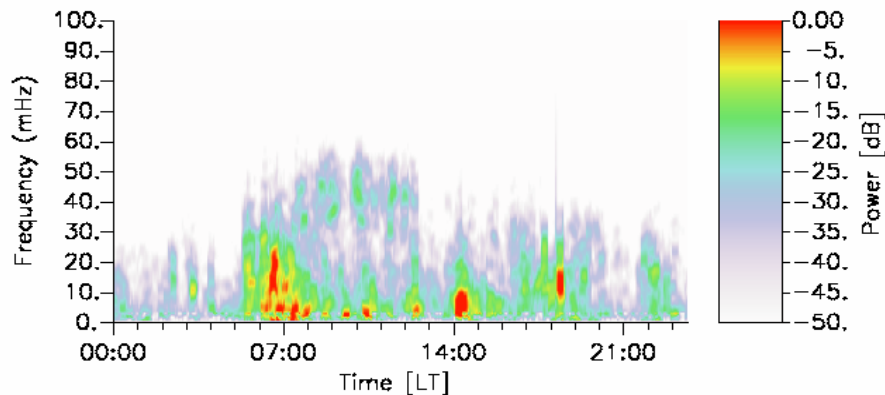
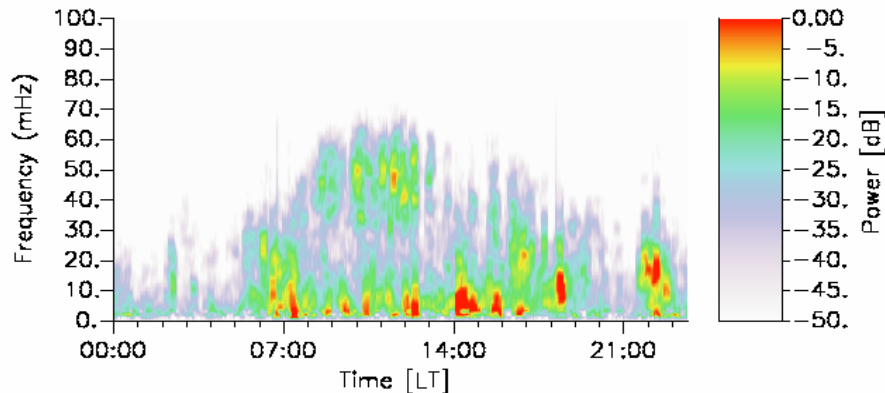


Example 1

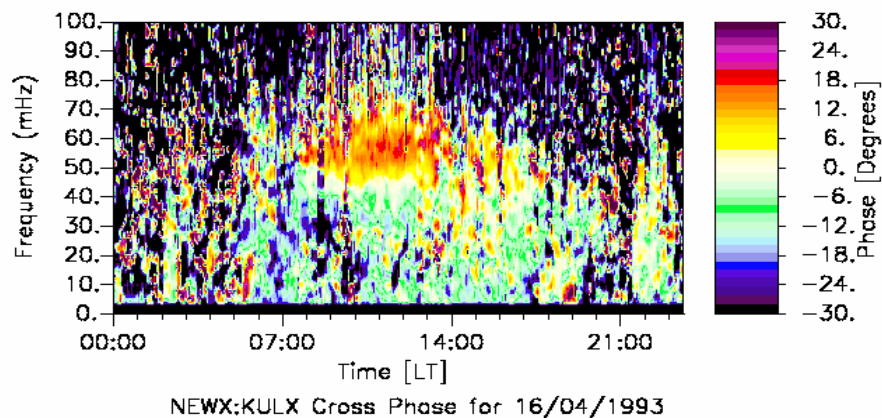
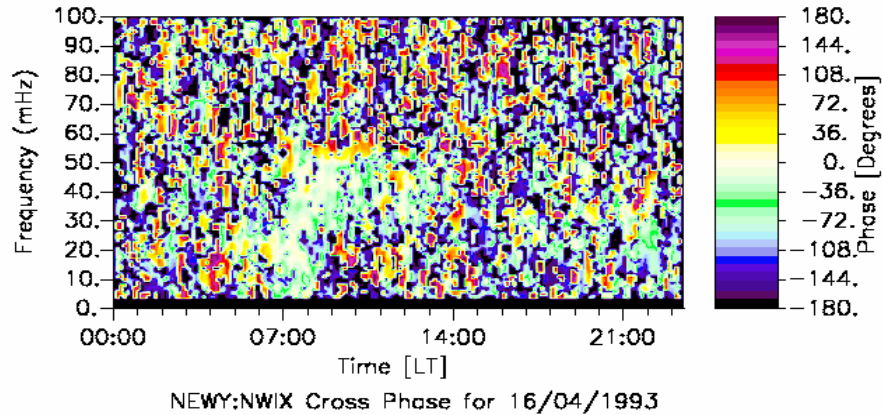
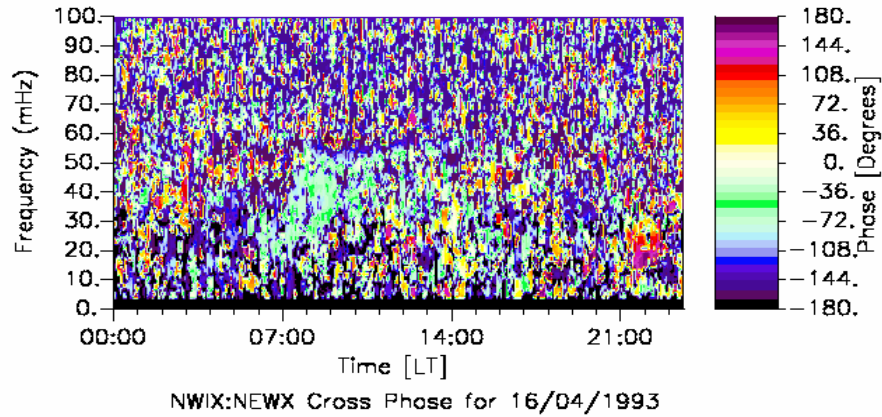


Example time series for L=1.9 magnetometer (upper panels) and co-located Doppler sounder, 16 April 1993.

Example 1



Whole-day power spectra for L=1.9 magnetometer (top two panels) and co-located Doppler sounder, 16 April 1993 (Kp=2 – 5).



Example 1

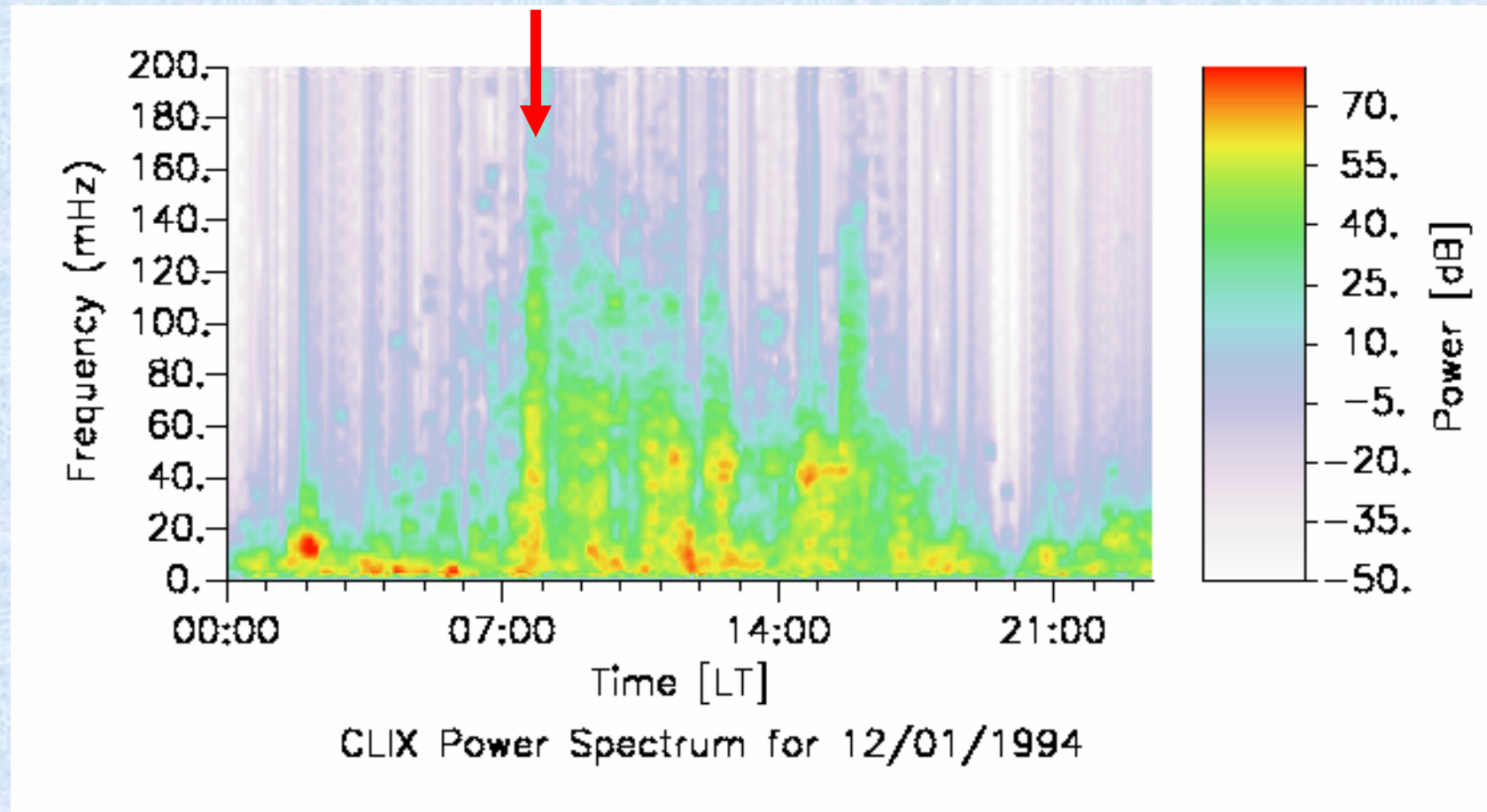
Whole-day cross-phase spectra at L=1.9 on 16 April 1993, between

(a) Doppler & NS mag

(b) Doppler & EW mag

(c) Closely spaced mags, NS component.

Example 2

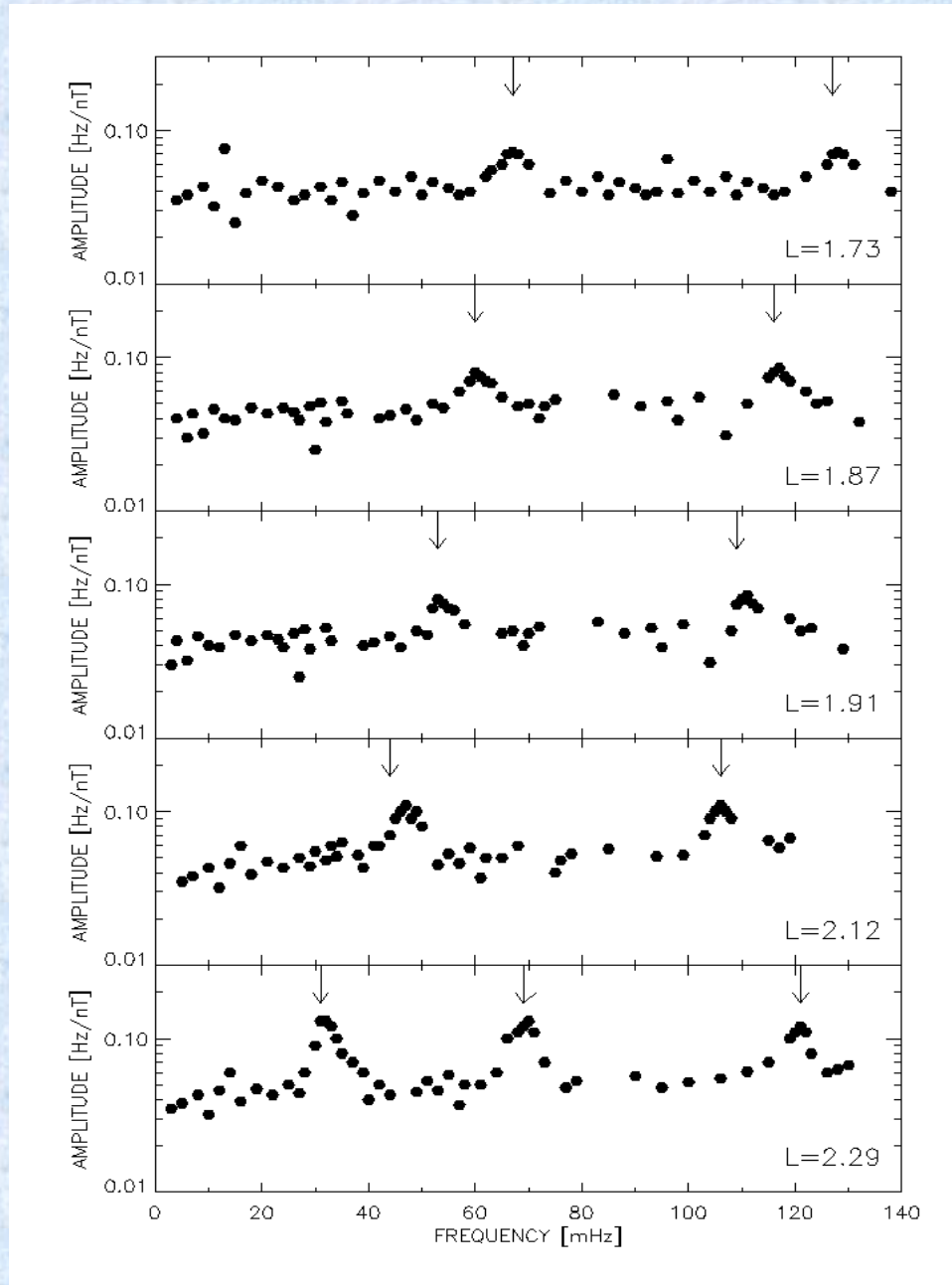


Pure state filtered Doppler power spectrum for 12 Jan 1994.

Example 2

Normalized amplitude of ionospheric Doppler oscillations as a function of frequency and latitude; 12 Jan 1994.

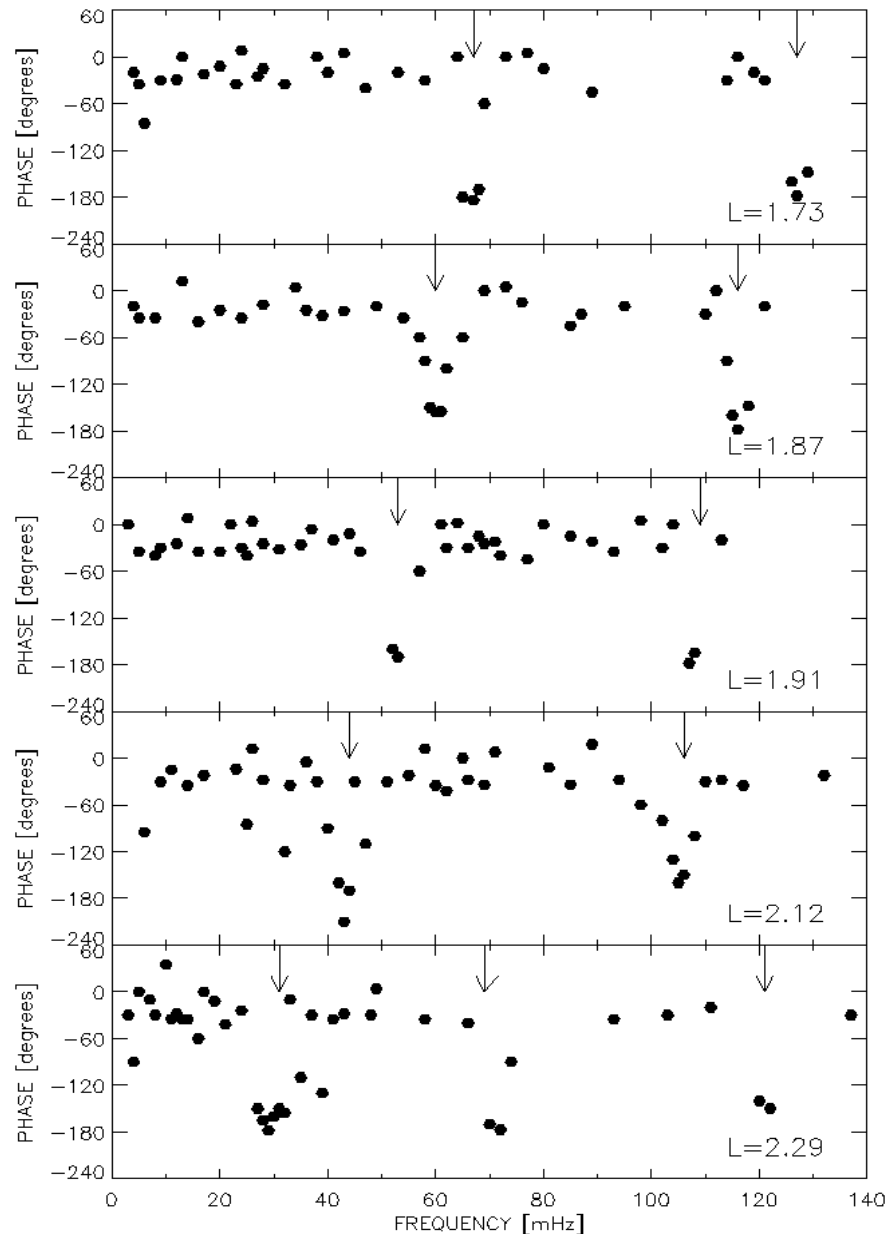
Arrows indicate local field line resonance.



Example 2

Ionosphere-ground phase of ULF Doppler oscillations as a function of frequency and latitude; 12 Jan 1994.

Arrows indicate local field line resonance.

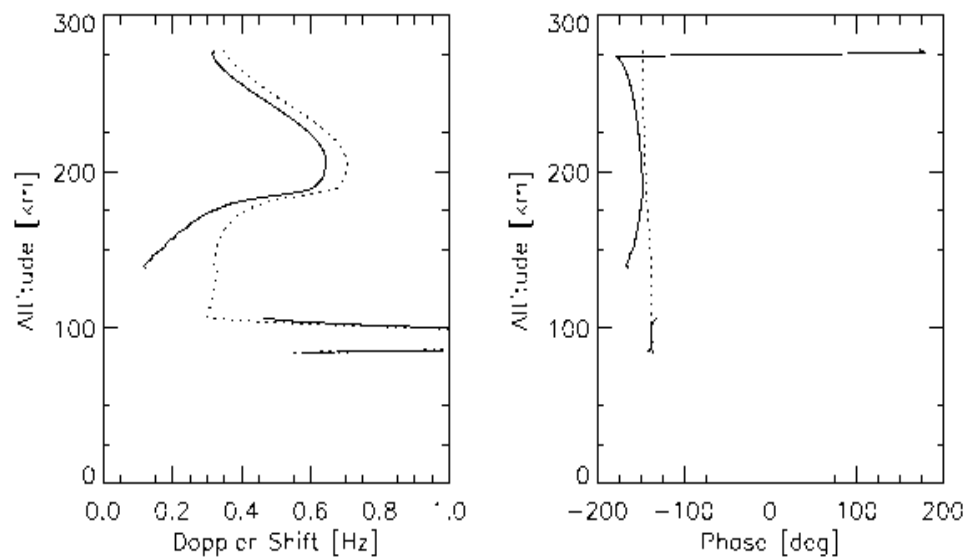
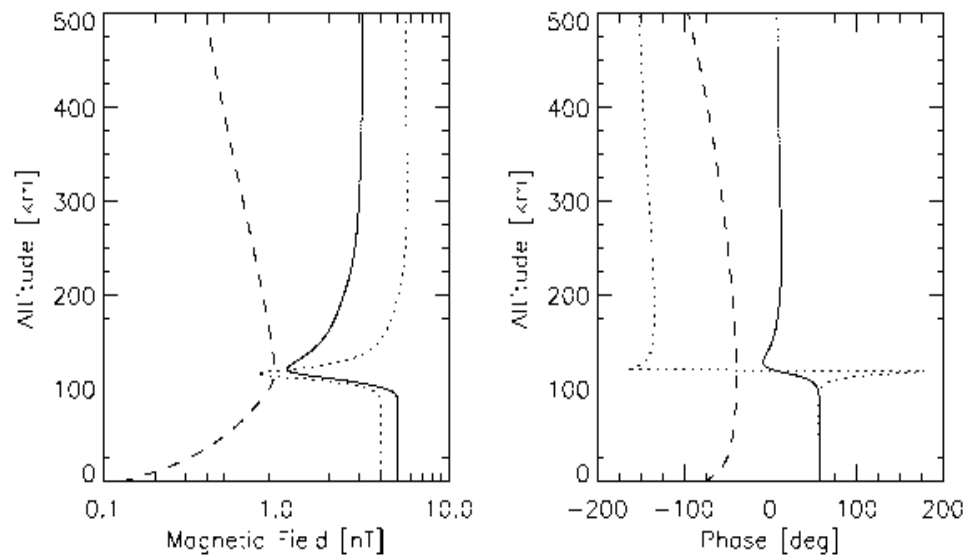


Modelling

- Used IRI-95 and MSIS-90 to represent ionospheric and atmospheric conditions, IGRF for the main magnetic field, and used measured B variations.
- Assumed incident ULF wave is 90 – 98% shear Alfvén mode at 53 mHz (resonance), but becomes 100% fast mode (compressional) at 43 and 63 mHz.
Set $k_y = 8 \times 10^{-7} \text{ m}^{-1}$ and $k_x = 1 \times 10^{-6} \text{ m}^{-1}$
- Then computed wave E and B altitude profiles, and Doppler shift of the radio signal due to changes in refractive index, advection, and compression by field-aligned component of B [Sutcliffe and Poole, *Planet. Space Sci.*, **38**, 1990; Sciffer et al., *Ann. Geophys.*, **23**, 2005].
- The main effect on ionospheric plasma is due to advection driven by electric field of downgoing ULF wave.

Model Results

B field and phase of downgoing ULF wave at 53 mHz, with 90% Alfvén mode, 10% fast mode. b_x =solid line, b_y =dots, b_z =dashes.



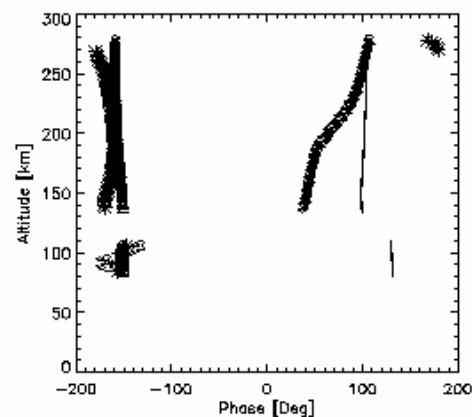
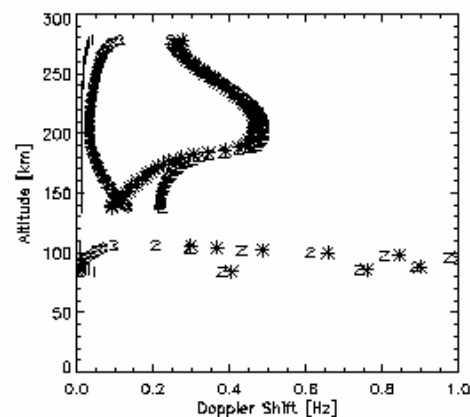
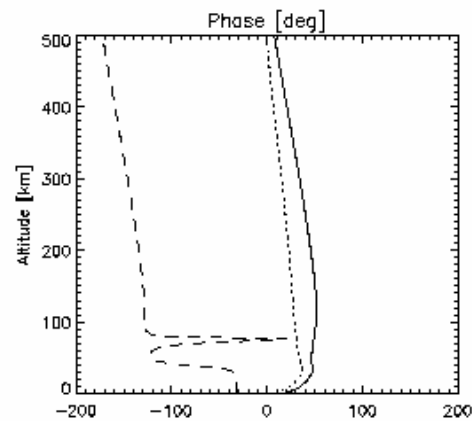
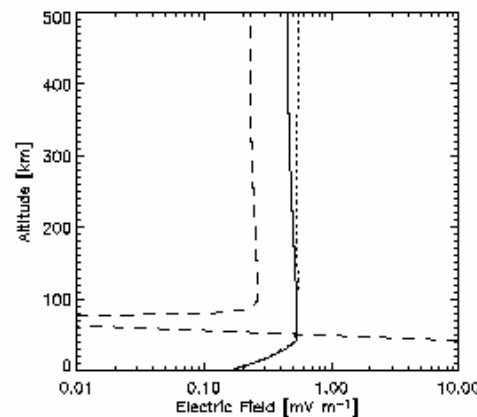
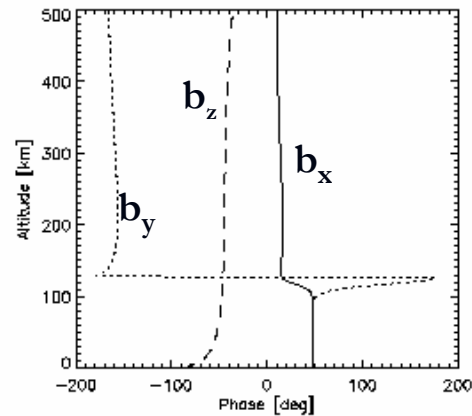
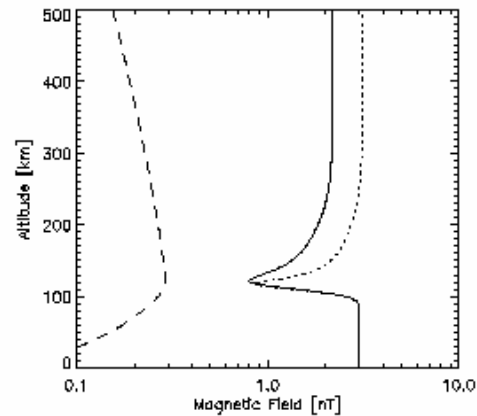
Resultant Doppler shift amplitude and phase. Solid line = total of all mechanisms; dotted line = vertical motion (advection) alone.

Previous Modelling

From Sutcliffe and Poole, *Planet. Space Sci.*, 38, 1581, 1990.

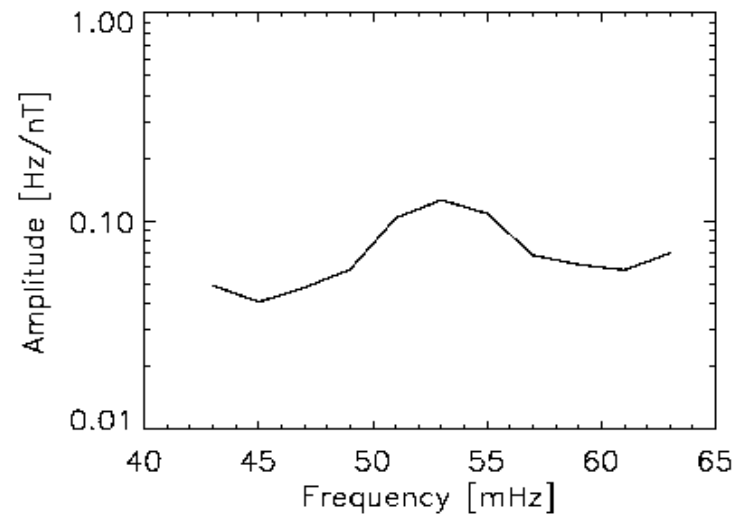
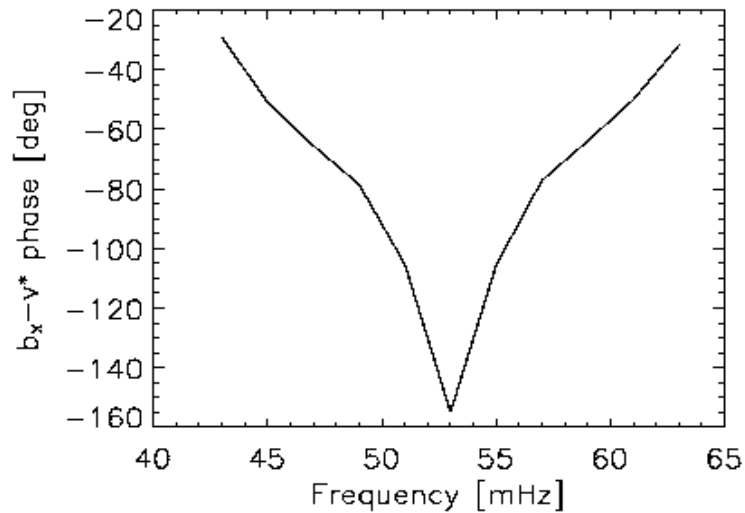
Amplitude and phase for shear Alfvén wave at 50 mHz (upper panels).

Resultant Doppler shifts (bottom) for V1, V2, V3 and total mechanisms



Model Results

Resultant ionosphere-ground phase (top) and amplitude ratio.

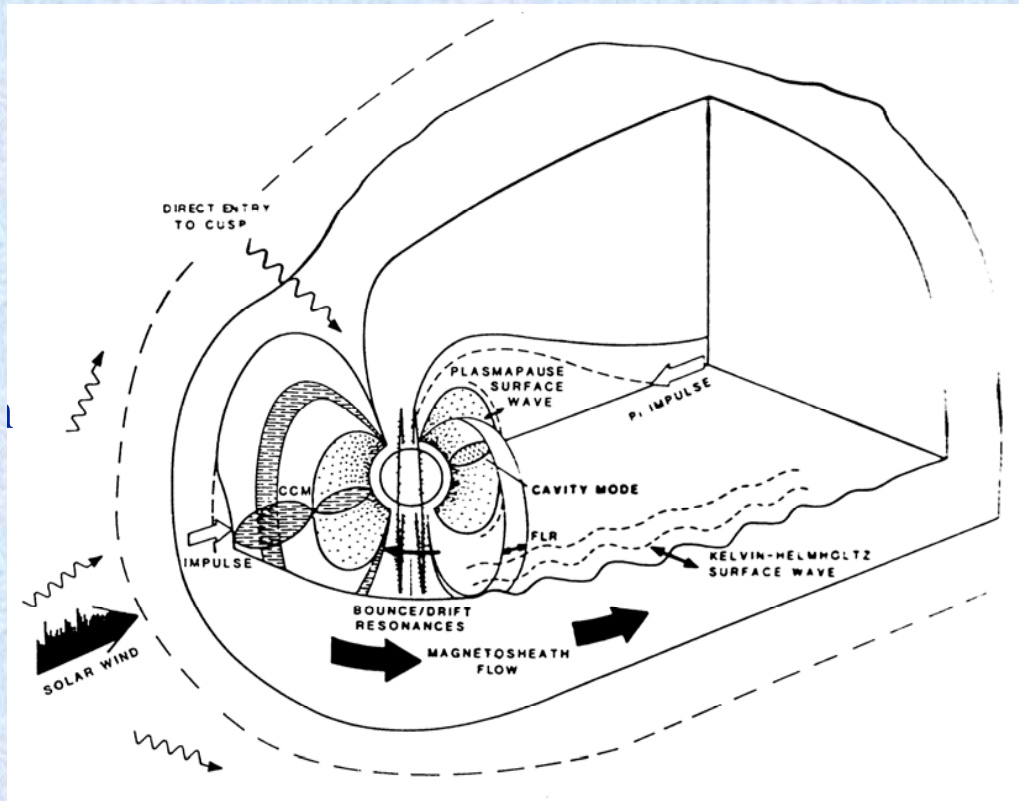


The incoming wave is assumed to comprise 98% Alfvén mode at 53 mHz, changing to 100% fast mode at 43 and 63 mHz.

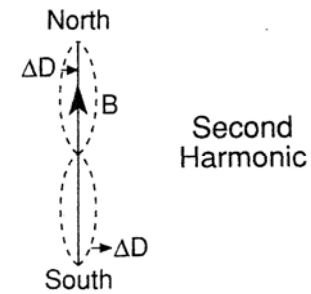
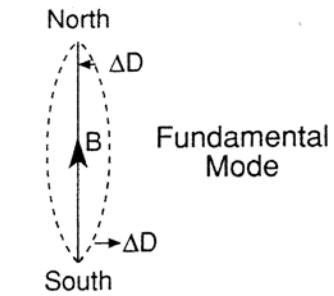
Conclusions

- In our low latitude study, almost every day when ionospheric Doppler oscillations were observable, they were synchronous with ULF waves recorded on the ground.
- Away from the resonant frequency the ionosphere-ground amplitude and phase are almost constant with frequency, around 0.05 Hz/nT and -30° .
- At the local resonant frequency there is a peak in the amplitude ratio and a pronounced dip in the phase. Further peaks and dips identify resonance harmonics.
- Modelling shows these features are related to the admixture of downgoing fast and shear Alfvén mode waves.

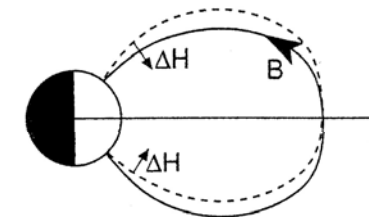
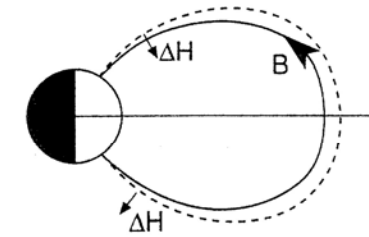
Incident Wave Modes



TORODIAL MODE



POLODIAL MODE



Instrumentation

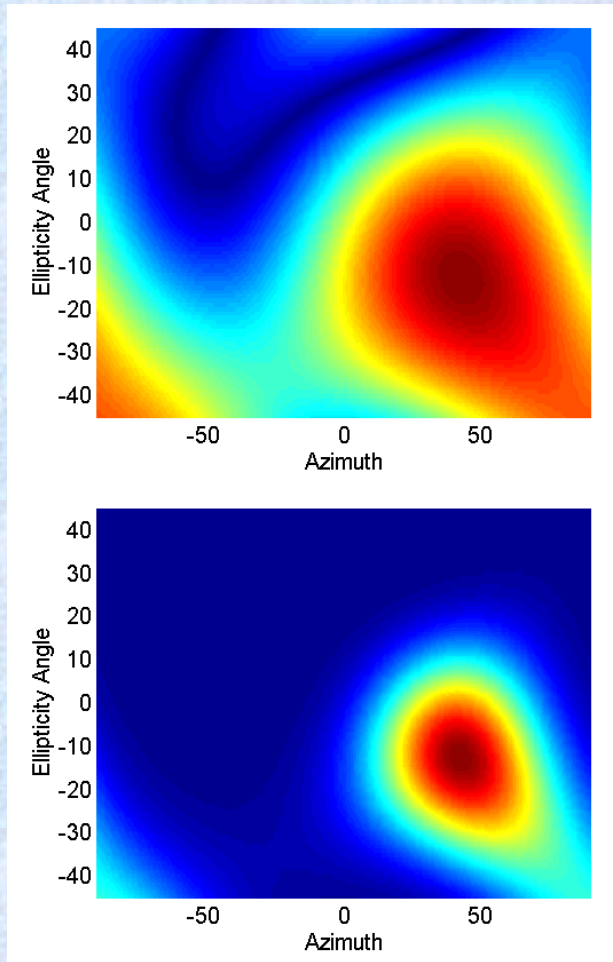
Magnetometers

- An array of up to 10 closely-spaced magnetometers recorded ULF pulsations at the ground.
- Cross-phase techniques were used to discriminate field line resonances from other wave modes.

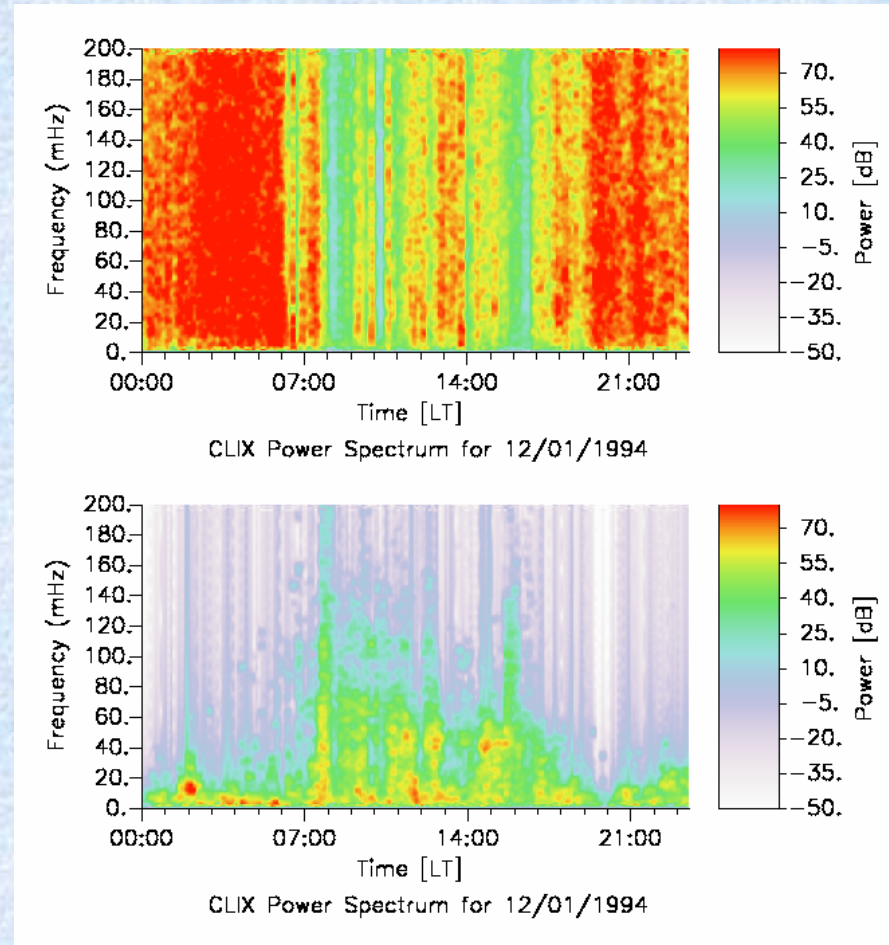
Doppler Sounders

- High stability HF transmitters at L=1.94 were monitored by an array of 5 HF Doppler receivers co-located with the magnetometers.
- Only the O-mode signal was recorded.
- Doppler sensitivity was ≤ 0.005 Hz over 1-140 mHz.

Doppler Analysis

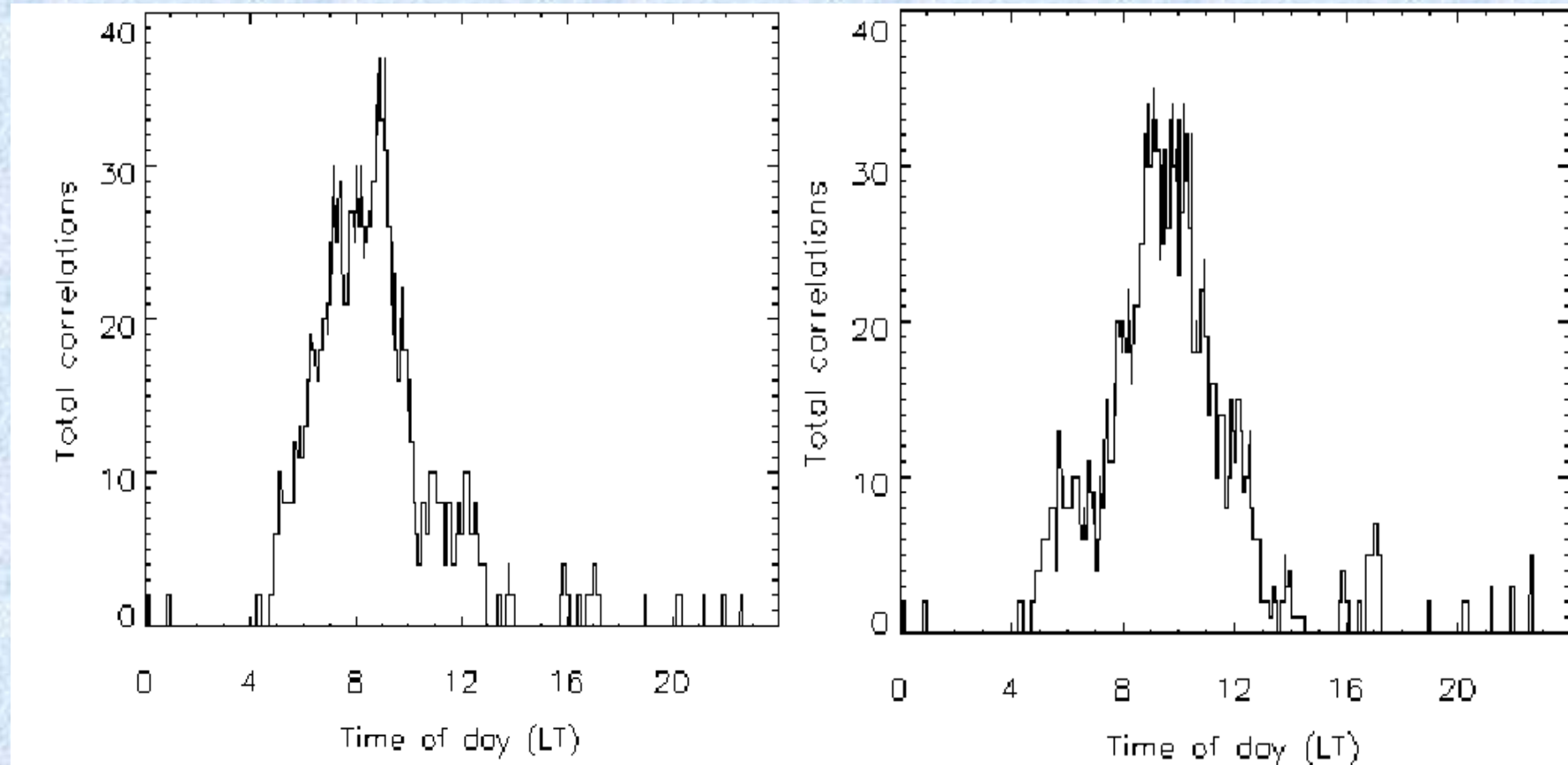


State vector response for Doppler data.



Unfiltered and filtered Doppler power spectra for 12 Jan 1994.

General Properties



Number of ULF events in the F-region and on ground NS (left) and EW (right) components over 32 days [30 min bins].